

Fermion propagation through spacetime

The electron propagator is given by

$$i S_{\alpha\beta}(x) \equiv \langle 0 | T[\psi_{\alpha}(x) \bar{\psi}_{\beta}(0)] | 0 \rangle$$

As we will see, have to define

$$T[\psi(x) \bar{\psi}(0)] = \theta(x^0) \psi(x) \bar{\psi}(0) - \theta(-x^0) \bar{\psi}(0) \psi(x)$$

↑
sign due to
anti-commuting nature
of ψ

Using

$$\psi(x) = \int \frac{d^3 p}{(2\pi)^{\frac{3}{2}} \left(\frac{E_p}{m}\right)^{\frac{1}{2}}} \sum_s [b(p,s) u(p,s) e^{-ip \cdot x} + d^{\dagger}(p,s) v(p,s) e^{ip \cdot x}] ,$$

we get for $x^0 > 0$

$$i S(x) = \langle 0 | \psi(x) \bar{\psi}(0) | 0 \rangle$$

$$\begin{aligned} &= \int \frac{d^3 p}{(2\pi)^{\frac{3}{2}} \left(\frac{E_p}{m}\right)^{\frac{1}{2}}} \int \frac{d^3 p'}{(2\pi)^{\frac{3}{2}} \left(\frac{E_{p'}}{m}\right)^{\frac{1}{2}}} \left[\langle 0 | \sum_{s,s'} u(p,s) \bar{u}(p',s') e^{-ip \cdot x} \underbrace{b b^{\dagger}}_{= \{b, b^{\dagger}\}} | 0 \rangle \right. \\ &\quad \left. + \langle 0 | \sum_{s,s'} u(p,s) \bar{v}(p',s') e^{-ip \cdot x} \underbrace{b d}_{= 0} | 0 \rangle \right] \\ &= \delta^{(3)}(\mathbf{p} - \mathbf{p}') \delta_{ss'} \end{aligned}$$

$$+ \underbrace{\langle 0 | d^\dagger}_{s, s'} \sum_{s, s'} u(p, s) \bar{u}(p', s') e^{ip \cdot x} b^\dagger | 0 \rangle$$

$$= (d | 0 \rangle)^\dagger = 0$$

$$+ \langle 0 | \sum_{s, s'} u(p, s) \bar{v}(p', s') e^{ip \cdot x} \underbrace{d^\dagger d}_{=0} | 0 \rangle$$

$$= \int \frac{d^3 p}{(2\pi)^3 E_{p/m}} \sum_s u(p, s) \bar{u}(p, s) e^{-ip \cdot x}$$

$$= \int \frac{d^3 p}{(2\pi)^3 E_{p/m}} \frac{\not{p} + m}{2m} e^{-ip \cdot x} \quad (1)$$

For $x^0 < 0$, we have to compute

$$iS_{\alpha\beta}(x) = - \langle 0 | \bar{\psi}_\beta(0) \psi_\alpha(x) | 0 \rangle$$

$$= - \int \frac{d^3 p}{(2\pi)^3 E_{p/m}} \sum_s \bar{v}_\beta(p, s) u_\alpha(p, s) e^{-ip \cdot x}$$

$$= - \int \frac{d^3 p}{(2\pi)^3 E_{p/m}} \left(\frac{\not{p} - m}{2m} \right)_{\alpha\beta} e^{-ip \cdot x} \quad (2)$$

Putting (1) and (2) together, we get

$$iS(x) = \int \frac{d^3 p}{(2\pi)^3 E_{p/m}} \left[\theta(x^0) \frac{\not{p} + m}{2m} e^{-ip \cdot x} - \theta(-x^0) \frac{\not{p} - m}{2m} e^{ip \cdot x} \right]$$

This can be written more elegantly as

$$\begin{aligned} iS(x) &= i \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot x} \frac{p+m}{p^2 - m^2 + i\varepsilon} \\ &= \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot x} \frac{i}{p - m + i\varepsilon} \end{aligned}$$

To see this, note that in the complex p^0 plane the integrand has poles at

$$p^0 = \pm \sqrt{\vec{p}^2 + m^2 - i\varepsilon} \simeq \pm (E_p - i\varepsilon)$$

For $x^0 > 0$, the factor $e^{-ip \cdot x}$ tells us to close the contour in the lower half-plane

→ going around the pole $+(E_p - i\varepsilon)$ clockwise, we get

$$iS(x) = (-i)i \int \frac{d^3 p}{(2\pi)^3} e^{-ip \cdot x} \frac{p+m}{2E_p}$$

For $x^0 < 0$, we have to close the contour in the upper half-plane

→ have to go around the pole $-(E_p - i\varepsilon)$ anticlockwise

$$\rightarrow iS(x) = i^2 \int \frac{d^3 p}{(2\pi)^3} e^{+iE_p x^0 + i\vec{p} \cdot \vec{x}} \\ \times \frac{1}{-2E_p} (-E_p \gamma^0 - \vec{p} \cdot \vec{\gamma} + m)$$

flipping \vec{p} we get

$$iS(x) = - \int \frac{d^3 p}{(2\pi)^3} e^{ip \cdot x} \frac{1}{2E_p} (E_p \gamma^0 - \vec{p} \cdot \vec{\gamma} - m) \\ = - \int \frac{d^3 p}{(2\pi)^3} e^{ip \cdot x} \frac{1}{2E_p} (\not{p} - m)$$

To summarize, we see that the momentum space fermion propagator has the elegant form

$$iS(p) = \frac{i}{\not{p} - m + i\epsilon}$$

$\rightarrow S(p)$ is the inverse of the Dirac operator, just as the scalar boson propagator

$$D(k) = \frac{1}{k^2 - m^2 + i\epsilon} \quad \text{is the inverse}$$

of the Klein-Gordon operator $k^2 - m^2$

§ 2.4 Grassmann Integrals and Feynman Diagrams for Fermions

Vacuum energy:

By definition, vacuum fluctuations occur when there are no sources to produce particles

→ for free scalar field theory we get

$$Z = \int \mathcal{D}\varphi e^{i \int d^4x \frac{1}{2} [\partial\varphi]^2 - m^2\varphi^2}$$
$$= C \left(\frac{1}{\det[\partial^2 + m^2]} \right)^{\frac{1}{2}} = C e^{-\frac{1}{2} \text{Tr} \log(\partial^2 + m^2)} \quad (1)$$

where we used the identity

$$\det M = e^{\text{Tr} \log M} \quad (2)$$

(exercise)

Recall that $Z = \langle 0 | e^{-iHT} | 0 \rangle$ (with $T \rightarrow \infty$)

$$= e^{-iET}$$

energy of vacuum

Let us now evaluate the trace in (1)

this can be done by noting

$$\begin{aligned}\text{Tr } G &= \int d^4x \langle x | G | x \rangle \\ &= \int d^4x \int \frac{d^4k}{(2\pi)^4} \int \frac{d^4q}{(2\pi)^4} \langle x | k \rangle \langle k | G | q \rangle \langle q | x \rangle\end{aligned}$$

→ we get :

$$ET = \frac{1}{2} VT \int \frac{d^4k}{(2\pi)^4} \log(k^2 - m^2 + i\epsilon) + A \quad (3)$$

where A is an infinite constant corresponding to the multiplicative factor C in (1)

Let us define m' by writing

$$A = -\frac{1}{2} VT \int \frac{d^4k}{(2\pi)^4} \log(k^2 - m'^2 + i\epsilon)$$

→ (3) gives difference of vacuum energies of particles with mass m and m'

$$\rightarrow \frac{E}{V} = -\frac{i}{2} \int \frac{d^4k}{(2\pi)^4} \log \left[\frac{k^2 - m^2 + i\epsilon}{k^2 - m'^2 + i\epsilon} \right]$$

$$= -\frac{i}{2} \int \frac{d^3k}{(2\pi)^3} \int \frac{d\omega}{2\pi} \log \left[\frac{\omega^2 - \omega_k^2 + i\varepsilon}{\omega^2 - \omega_k'^2 + i\varepsilon} \right]$$

where $\omega_k' = +\sqrt{k^2 + m^2}$

Integrating by parts, we get

$$\begin{aligned} & \int \frac{d\omega}{2\pi} \log \left[\frac{\omega^2 - \omega_k^2 + i\varepsilon}{\omega^2 - \omega_k'^2 + i\varepsilon} \right] \\ &= -2 \int \frac{d\omega}{2\pi} \omega \left[\frac{\omega}{\omega^2 - \omega_k^2 + i\varepsilon} - \frac{\omega}{\omega^2 - \omega_k'^2 + i\varepsilon} \right] \\ &= -i2\omega_k^2 \left(\frac{1}{-2\omega_k} \right) - (\omega_k \rightarrow \omega_k') \\ &= +i(\omega_k - \omega_k') \end{aligned}$$

Restoring t_1 , we then arrive at

$$\frac{E}{V} = \int \frac{d^3k}{(2\pi)^3} \left(\frac{1}{2} t_1 \omega_k - \frac{1}{2} t_1 \omega_k' \right) \quad (4)$$

→ this matches with the result previously obtained using the operator formalism (see §1.1)

A peculiar sign for fermions

Recall that for fermions, the operator formalism gave a different sign (see § 2.3)

Question: How can we obtain this from the path integral formalism?

Recall that the origin of the sign lies in Gaussian integration result

$$\int_{-\infty}^{\infty} dx e^{-\frac{1}{2}ax^2} = \sqrt{\frac{2\pi}{a}} = \sqrt{2\pi} e^{-\frac{1}{2}\log a}$$

→ need to find new type of integral giving $e^{+\frac{1}{2}\log a}$

Grassmann variables

We postulate, following Grassmann (19th century), the existence of a new kind of number, called Grassmann or anticommuting number:

$$\begin{aligned} \eta \zeta &= -\zeta \eta \\ \rightarrow \eta^2 &= 0 \end{aligned}$$

functions of η are of general form:

$$f(\eta) = a + b\eta$$

postulate translation inv. of integral:

$$\int_{-\infty}^{\infty} d\eta f(\eta + \zeta) = \int_{-\infty}^{\infty} d\eta f(\eta)$$

$$= \int_{-\infty}^{\infty} d\eta (a + b(\eta + \zeta)) = \int_{-\infty}^{\infty} d\eta (a + b\eta)$$

$$\rightarrow \int_{-\infty}^{\infty} d\eta b\zeta = 0 \quad \text{for any Grassmann number } \zeta$$

thus we conclude: $\int d\eta b = 0 \rightarrow \int d\eta 1 = 0$
ordinary number

Since $\int d\eta \eta$ behaves like ordinary number, we define $\int d\eta \eta = 1$